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# E1 SIGNAL PROCESSING OF THE GALILEO SYSTEM IN THE **NAVIGATION RECEIVER**

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### Resume

The subject of this article are issues related to the navigation system in the field of analyzing the processed signal in the GNSS system receiver. The main purpose of the work is to discuss the Galileo E1 signal processing methods in the GNSS navigation system receiver, supported by adapted research tools in terms of solving the research problem (analysis, model, simulation tests) and the mathematical apparatus used. Key studies are concentrated around the process of generating the navigation data, dispersing sequences and signal modulation. Thus, when designing a receiver, it is better to use the simulation signals than the real ones, since one can get more control over the properties of the received signal. In the final part of the work, in accordance with the subject of research, based on the developed appropriate research tools, observations and final conclusions were formulated, which have practical applications.

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### Introduction

In recent years, many efforts have been made to design and implement the Galileo signal simulator. This work presents simulations of the GNSS/Galileo navigation system transmitter signals via a graphical programming language, which is the Matlab/Simulink programming environment, with the same efficiency as text programming languages.

The GUI (Graphical User Interface) environment was used. In addition, use of a graphical programming language makes every part of the navigation system transmitter architecture very clear and much easier to understand and modify. The Galileo system receiver consists of four main components: acquisition, satellite tracking, digital processing and positioning efficiency. It is based entirely on the algorithms of the Matlab/ Simulink program [1-2].

In addition, some of the decoding algorithms were previously developed for the GPS system technology, so everything had to be adjusted, including the navigation message. However, it should be noted that majority of the Galileo system algorithms have been recently developed.

Real Galileo signals taken from the GNSS navigation system receiver were used for the tests. These signals were helpful, among others in understanding the structure of navigational messages and obtaining parameters transmitted by the Galileo system signals. However, it should be mentioned that study encountered some inconvenience due to insufficient information on the system under consideration. Although there is a lot of information and publications in the field of the GPS, the Galileo system is still under development and some specifications have not been fully defined or have been left to ongoing improvement.

In the rest of this article, in accordance with the subject of the study, the issue of the positioning efficiency of the GNSS navigation system (Galileo) was considered from an aspect of an ideal and not ideal mathematical model of the pseudo-distance of the system [3-5].

### $\mathbf{2}$ Positioning performance

Navigation data consists of 4 types of data needed to perform positioning, i.e. ephemeris parameters necessary to indicate the position of the satellite relative to the receiver, time and clock correction parameters to calculate the pseudorange, service parameters for identifying satellites and signal quality and almanac parameters, indicating the positions of other satellites. After presenting the ephemeris, services and almanac parameters, it is time to delve into time parameters and clock correction. This is necessary to calculate the so-called pseudorange. Then the equations used to find the position of the satellite would be presented.

### 2.1 Pseudorange of the ideal model

The relative pseudorange is the distance (or time) between the two reference points. There is no absolute reference to time in the GNSS system. The only reference to time is the sampling frequency and the clock deviation of the receiver [6-8].

To determine its position, the receiver determines the distances from at least four satellites, as well as their positions at the time of broadcasting. Knowing the orbital parameters of the satellites, these positions can be calculated for any point in time. The pseudorange of each satellite is obtained by multiplying the speed of light by the time the signal traveled from the satellite to the receiver [9-11]. The ideal case for calculating the pseudorange is when there are no errors with the transmitter and receiver clock, ionosphere, troposphere or receiver noise. This case is calculated according to:

$$R_{i}(t_{T},t_{R}) = |r_{s_{i}}(t_{T_{i}}) - r_{R}(t_{R})| = = \sqrt{(x_{s_{i}} - x_{R})^{2} + (y_{s_{i}} - y_{R})^{2} + (z_{s_{i}} - z_{R})^{2}},$$
(1)

where:  $R_i(t_T, t_R)$  - the distance between the transmitter and the receiver,  $t_R$  - the moment of time when the received signal arrives at the receiver,  $t_{T_i}$  - the moment of time when the signal is sent by satellite i.

One can also define the delay associated with the satellite i, marked  $\tau_i$  and written as:

$$R_{i}(t_{T},t_{R}) = |r_{s_{i}}(t_{T_{i}}) - r_{R}(t_{R})|\tau_{i} = \frac{R_{i}(t_{T},t_{R})}{c} = t_{R} - t_{T_{i}}.$$
(2)

At least 3 satellites are required to calculate position. A non-linear system of three equations with three unknowns must be solved.

### 2.2 Pseudorange of the non-ideal model

This is the real case where errors occur. The transmission time of the satellite i is:

$$t_{T_i} = t_{T_i}^{Galileo} + \Delta t_{T_i}, \qquad (3)$$

where:  $\Delta t_{Ti}$  - satellite clock error.

The signals reach the receiver at the following moment, it can be written as:

$$t_R^{Galileo} = t_{T_i}^{Galileo} + \tau_i - \Delta t_1^{rel}$$
, (4)

where:  $\Delta t_i^{rel}$  - relativistic correction.

However, the moment measured by the receiver, taking into account the above Equation (4), can be presented in the following form:

$$t_R = t_R^{Galileo} + \Delta t_R + \Delta t_n = t_{T_i}^{Galileo} + \tau_i - \Delta t_i^{rel} + \Delta t_R + \Delta t_n,$$
(5)

where:  $\Delta t_R$  - receiver clock error,  $\Delta t_n$  - delay caused by the receiver measurement error due to the white noise. In contrast, the random variable has the same statistical properties for each satellite. Finally, taking into account Equation (3), the following form was obtained [12-14]:

$$t_R = t_{T_i} + \tau_i - (\Delta t_{T_i} + \Delta t_i^{rel}) + \Delta t_R + \Delta t_n.$$
 (6)

It should be noted that  $\Delta t_R$  is common to all satellites. The delay  $\tau_i$  is obtained by the sum of the geometric  $\tau_i^{geo}$ , ionospheric  $\Delta t_i^{ion}$  and tropospheric  $\Delta t_i^{trop}$  delays, as shown in the following form:

$$\tau_i = \tau_i^{geo} + \Delta t_i^{ion} + t_i^{trop}. \tag{7}$$

The total delay between transmitter and receiver is:

$$\boldsymbol{\tau}_{i}^{tot} = t_{Ri} - t_{Ti} = \boldsymbol{\tau}_{i} - (\Delta t_{Ti} + \Delta t_{i}^{rel}) + \Delta t_{R} + \Delta t_{n}. \quad (8)$$

Replacing  $\tau_i$  by Equation (7), the following form was obtained:

$$\tau_i^{tot} = \tau_i^{geo} + \Delta t_i^{ion} + \Delta t_i^{trop} - (\Delta t_{T_i} + \Delta t_i^{rel}) + \Delta t_R + \Delta t_n$$
(9)

After calculating the total delay, it is possible to obtain a pseudorange by multiplying it by the speed of light c, resulting in [15-17]:

$$\rho_i = c\tau_i^{tot} = R_i + c\Delta t_i^{ton} + c\Delta t_i^{trop} - \\
- c(\Delta t_{T_i} + \Delta t_i^{rel}) + c\Delta t_R + c\Delta t_n. \tag{10}$$

where:  $R_i$  - the geometric distance and as already shown in Equation (1) of the ideal model, the geometric distance is calculated as follows:

$$R_i = R_i(t_{T_i}^{Galileo}, t_R^{Galileo}) = |r_{si}(t_{T_i}^{Galileo}) - r_R(t_R^{Galileo})|,$$
 (11)

$$R_{i}(t_{T_{i}}^{Gal}, t_{R}^{Gal}) = \sqrt{\frac{\left[x_{s_{i}}(t_{T_{i}}^{Gal}) - x_{r}(t_{R}^{Gal})\right]^{2} + \left[y_{s_{i}}(t_{T_{i}}^{Gal}) - y_{r}(t_{R}^{Gal})\right]^{2} + \dots + \left[z_{s_{i}}(t_{T_{i}}^{Gal}) - z_{r}(t_{R}^{Gal})\right]^{2}}$$
(12)

It should be noted that both the parameter  $\Delta t_R$  from Equation (10) and the parameter  $r_R$  from Equation (11) do not depend on the position of the satellites. The parameter  $\Delta t_R$ , i.e. the receiver clock error, is not known, so it can be treated as another unknown. Hence, four satellites are necessary to determine coordinates of the receiver and its clock error.

## 3 Linearization

To solve the obtained system of equations, it should be taken into account that it is non-linear. It should be linearized, assuming that the approximate position and E48 SETLAK, KOWALIK

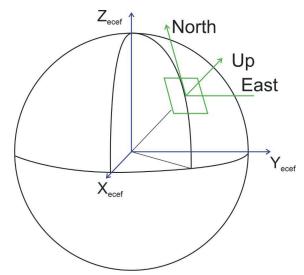


Figure 1 Coordinate system used in the mathematical description

approximate clock error are known. The pseudorange is derived from the following formula [18-20]:

$$\delta 
ho_i = \frac{\delta 
ho_i}{\delta x_R} \delta x_R + \frac{\delta 
ho_i}{\delta y_R} \delta y_R + \frac{\delta 
ho_i}{\delta z_R} \delta z_R + \frac{\delta 
ho_i}{\delta w_R} \delta w_R$$
,(13)

where:  $w_R = c\Delta t_R$ .

In turn, partial derivatives can be written in the

$$\frac{\delta \rho_i}{\delta x_R} = \frac{-x_{S_i} - x_R}{R_i},\tag{14}$$

$$\frac{\delta \rho_i}{\delta v_P} = \frac{-y_{S_i} - y_R}{R_i},\tag{15}$$

$$\frac{\delta \rho_i}{\delta z_P} = \frac{-z_{S_i} - z_R}{R_i},\tag{16}$$

$$\frac{\delta \rho_i}{\delta w_R} = 1. \tag{17}$$

Then it is defined as:

$$\delta_{p} = [\delta x_{R} \delta y_{R} \delta z_{R} \delta w_{R}]^{T}; \delta_{p} = [\delta \rho_{1} \delta \rho_{2}...\delta \rho_{K}]^{T}, \quad (18)$$

$$A = \begin{pmatrix} a_{11} & a_{12} & a_{13} & 1 \\ a_{21} & a_{22} & a_{23} & 1 \\ \vdots & \vdots & \ddots & \vdots \\ a_{K1} & a_{K2} & a_{K3} & 1 \end{pmatrix}, \tag{19}$$

where: 
$$a_{i1} = \frac{-x_{si} - x_R}{R_i}, a_{i2} = \frac{-y_{si} - y_R}{R_i}, a_{i3} = \frac{-z_{si} - z_R}{R_i}$$
 and  $p = [x_R y_R z_R w_R]^T$  - specifies the extended vector

with receiver error.

It can therefore be written in the following form:

$$\delta \rho = A \delta p$$
. (20)

The position of the receiver can be calculated iteratively. Regardless of whether an approximate extended position vector is available at the moment, the estimation can be improved as follows:

$$p^{(k)} = p^{(k-1)} + \delta p^{(k)}, \qquad (21)$$

where:  $\delta p^{(k)}$  - solution of the linear system.

However, due to the fact that it is possible to use more equations than unknown quantities, the solution will be the so-called the least squares method [21-23]:

$$\delta p^{(k)} = (A^{(k-1)T} A^{(k-1)})^{-1} A^{(k-1)T} \delta \rho^{(k-1)}, \qquad (22)$$

where:  $\delta 
ho^{(k)} = \hat{
ho} - 
ho^k$ .

This process will continue until the condition is met  $\delta p < 10^{-12}$ 

### 3.1 Calculation of the satellite position

To obtain the position of the satellite in the Earthrelated coordinate system ECEF (Earth-Centered Earth-Fixed), the issues considered in this section use all information obtained from ephemeris data.

The calculated average movement  $n_0$ , is the average angular velocity of the satellite:

$$n_0 = \frac{2\pi}{T} = \sqrt{\frac{\mu}{A^s}}$$
, (23) where:  $\mu$  - geocentric gravitational constant,  $A$  - semi-

After obtaining the average movement in  $\left[\frac{rad}{s}\right]$ , the corrected average movement can be calculated

$$n = n_0 + \Delta n, \qquad (24)$$

where:  $\Delta n$  - average movement difference from the calculated value.

The actual total time difference between time t and epoch  $t_{0s}$  is  $t_{b}$ , which can be determined by the following:

$$t_k = t - t_{0_s},$$
 (25)

where: *t* - is the Galileo system time.

The average anomaly can be determined from the following:

$$M = M_0 + nt_k, (26)$$

where:  $M_0$  - mean anomaly at reference time.

The Kepler equation for the eccentric E anomaly (can be solved by iteration) is defined as:

$$M = E - e\sin(E), (27)$$

where: e - eccentricity.

The condition of relativistic correction can be obtained by:

$$\Delta t_r = FeA^{\frac{1}{2}}\sin(E), \qquad (28)$$

where:  $F = -2\mu^{\frac{1}{2}}/c^2$  - constant.

The satellite time correction (in seconds) is modeled by the following second order polynomial that defines parabola according to the following equation [27-29]:

$$\Delta t_{SV}(X) = a_{f0} + a_{f1}(X)[t - t_{0C}(X)] + a_{f2}(X)[t - t_{0C}(X)]^2 + \Delta t_r,$$
(29)

where:  $a_{f0}, a_{f1}, a_{f2}$  - SV (Space Vehicle) clock correction coefficients,  $t_{0C}$  - clock correction related to Time of Week,  $(X) = (f_1, f_2)$  - frequency combination  $f_1$  and  $f_1$  used in the clock model.

The true anomaly is defined as:

$$v = \tan^{-1} \left\{ \frac{\sin(v)}{\cos(v)} \right\} = \tan^{-1} \times \left\{ \frac{\sqrt{1 - e^2} \sin(E) / (1 - e \cos(E))}{(\cos(E) - e) / (1 - e \cos(E))} \right\}.$$
 (30)

The argument of latitude can be determined from the formula:

$$u = v + \omega$$
. (31)

The argument of latitude correction is calculated as follows:

$$\delta u = C_{us} \sin(2\Phi) + C_{uc} \cos(2\Phi). \tag{32}$$

The radius vector correction is determined from the following formula:

$$\delta r = C_{rs}\sin(2\Phi) + C_{rc}\cos(2\Phi). \tag{33}$$

However, the inclination correction is determined from the following equation:

$$\delta i = C_{is}\sin(2\Phi) + C_{ic}\cos(2\Phi). \tag{34}$$

The anomaly at epoch t is the sum of the argument of latitude and its correction:

$$u = \Phi + \delta u . \tag{35}$$

The corrected radius vector is calculated from the formula:

$$r = A(1 - e\cos(E)) + \delta r. \tag{36}$$

In contrast, improved inclination is:

$$i = i_0 + \delta i + {\binom{\bullet}{i}} t_k. \tag{37}$$

Then the position in the orbital plane can calculated

$$x' = r\cos u, y' = r\sin u. \tag{38}$$

The right-ascension of the ascending node  $\Omega$  is calculated as follows according to equation [30-32]:

$$\Omega = \Omega_0 + {\binom{\bullet}{\Omega}} t_k. \tag{39}$$

However, this result is obtained for the geocentric inertial coordinate system ECIS (Earth-Centered Inertial System). After receiving  $\Omega$ , the next step to get the result in the ECEF system is:

$$\Omega = \Omega - \omega_E t_k - \omega_E t_{0s}, \tag{40}$$

where:  $\omega_E$  - average angular velocity of the Earth.

Finally, to find the satellite position, the following form was obtained:

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} x' \cos(\Omega) - y' \sin(\Omega)\cos(i) \\ x' \sin(\Omega) + y' \cos(\Omega)\cos(i) \\ y' \sin(i) \end{pmatrix}.$$
 (41)

### 3.2 BOC modulation

The Galileo system signals are modulated using a new technique called BOC (Binary Offset Carrier) modulation. The BOC modulation was designed to modernize the GPS and Galileo systems to facilitate addition of other signals that have the same carrier frequency in the radio frequency bands, without interfering with other signals.

Advantages of the BOC modulation are: improvement of traditional GNSS system signal properties for better resistance to multi-path and interference with various types of noise and ensuring the spectral isolation between signals with the same carrier frequency.

The BOC modulated signal consists of a sinusoidal carrier, a rectangular subcarrier (*BOC*sin and *BOC*cos), a pseudo-random PRN code (Pseudorandom Noise) and a data sequence. The signal is presented as BOC

E50

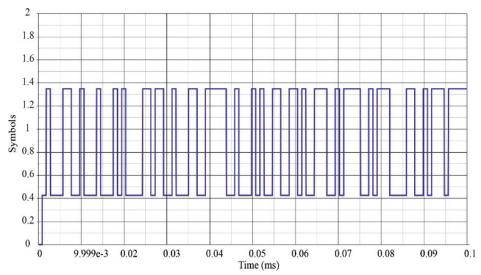


Figure 2 CBOC subcarrier of B channel with BOC (1.1) and BOC (6.1)

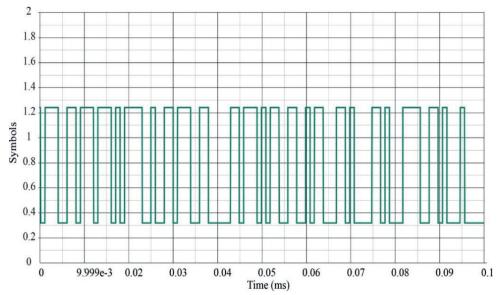
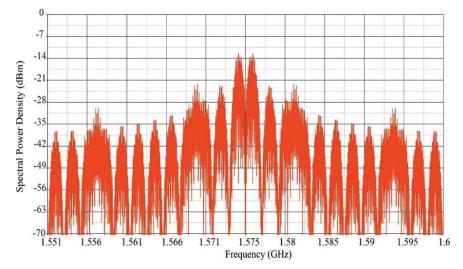


Figure 3 CBOC subcarrier of C channel with BOC (1.1) and BOC (6.1)



 $\textbf{\textit{Figure 4} Spectral power density of E1 signal for the CBOC modulation (1.6)}$ 

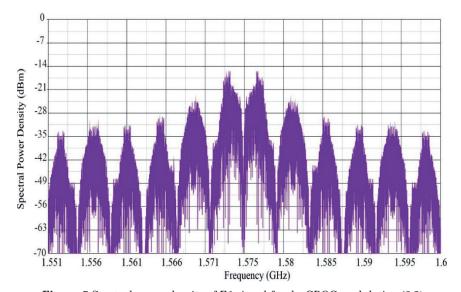


Figure 5 Spectral power density of E1 signal for the CBOC modulation (6.2)

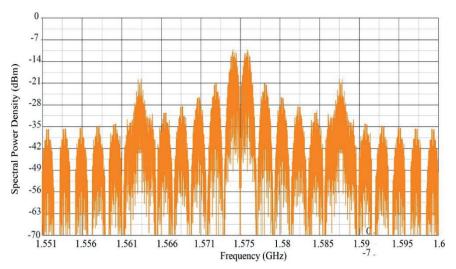


Figure 6 Spectral power density of E1 signal for the CBOC modulation (12.1)

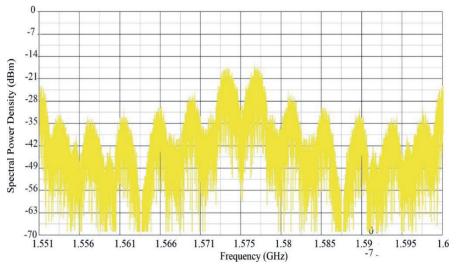


Figure 7 Spectral power density of E1 signal for the CBOC modulation (2.20)

E52

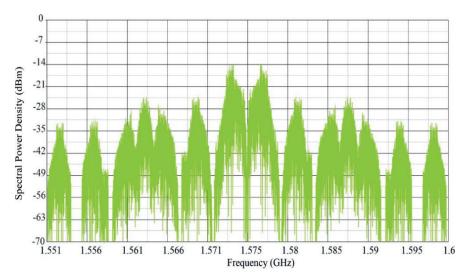
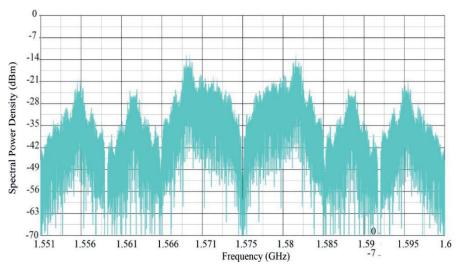


Figure 8 Spectral power density of E1 signal for the CBOC modulation (2.12)



 $\textbf{\textit{Figure 9}} \ Spectral \ power \ density \ of \ E1 \ signal \ for \ the \ CBOC \ modulation \ (6.12)$ 

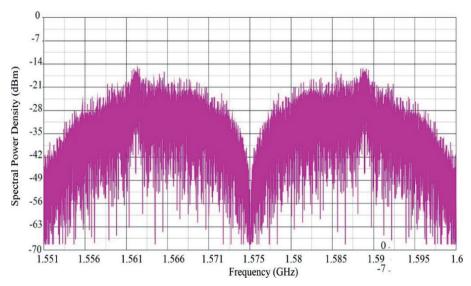


Figure 10 Spectral power density of E1 signal for the CBOC modulation (12.12)

(m,n). Two independent parameters m and n are used to concentrate the signal strength in specific parts of the spectrum to reduce interference with other signals [33-34].

Parameters m and n are defined as follows:

$$m = \frac{F_s}{F_0}, n = \frac{F_c}{F_0}, \tag{42}$$

where:  $F_s$  - defines the subcarrier frequency in [MHz],  $F_c$  - means transmission speed in [Mcps],  $F_o$  - reference frequency  $F_o$  = 1.023 [MHz] which is generated by the atomic clock on satellites.

### 4 Selected tests for the BOC modulation

Channels B and C are modulated by "Composite BOC" modulation (CBOC (6.1, 1/11)). CBOC modulation combines BOC (1.1) and BOC (6.1).

The CBOC modulation results from the multiplexing of the narrowband BOC signal (1.1), whose subcarrier frequency is 1.023 [MHz] and the transmission speed 1.023 [Mcps], with the wideband BOC signal (6.1), with the subcarrier frequency 6.138 [MHz] and the transmission speed 1.023 [Mcps].

The normalized CBOC power spectral density (6.1, 1/11) is obtained according to equation:

$$G_{CBOC}(6.1\frac{1}{11})(f) = \frac{10}{11}G_{BOC}(1.1) + G_{BOC}(6.1)(f),$$
(43)

where:  $G_{CBOC(m,n)}(f)$  - PSD (Packet-Switched Data) unit of BOC sinusoidal modulation.

Thus 1/11 of the power is allocated on the high frequency channel (BOC (6.1)). CBOC modulation for B and C channels can be illustrated by the following equations:

$$scB(t) = \alpha sc(t)_{BOC(1.1)} + \beta sc(t)_{BOC(6.1)},$$
 (44)

$$scC(t) = \alpha sc(t)_{BOC(1,1)} + \beta sc(t)_{BOC(6,1)},$$
 (45)

where:  $\alpha=\sqrt{\frac{1}{10}}$  and  $\beta=\sqrt{\frac{10}{11}}$ . Figures 2 and 3 show the generated CBOC

Figures 2 and 3 show the generated CBOC subcarriers for scB(t) and scC(t) using the Matlab/Simulink environment.

Equations (46) and (47) identify *BOC*sin and *BOC*cos respectively:

$$BOC\sin sc(t) = sign(\sin(2\pi Fst)),$$
 (46)

$$BOC\cos sc(t) = sign(\cos(2\pi Fst)),$$
 (47)

where:  $F_s$  - subcarrier frequency.

It should be noted that the BOC modulation divides the spectrum into two main symmetrical layers focused on  $\pm f_s$  [MHz] (subcarrier frequency) around the carrier

frequency. The generated PSD from the SE1 signal is shown in Figures 4-10.

Figures 4-10, illustrate the signal spectra modulated by the composite CBOC method (n,m), differing in the occurrence of the main leaves representing the main band in which the transmission of the Galileo system radio signals is carried out and the shift, causing the use of additional band for other signal transmissions or services.

The above process is possible both by changing the transmission speed of the carrier wave n and the transmission speed of m chips. Based on the analysis of the spectra of the signal modulated on individual waveforms, it can be observed how the process of extending, adding a new band or its shift takes place for different values of n and m.

Especially valuable in this context is the signal power spectral density analysis E1 for the CBOC modulation (6.12) as shown in Figure 9, considering a candidate/pretender signal providing a wide frequency band E5a/E5b for the transmission of Galileo navigation signals.

It should be noted that the CBOC (6.12) signal is very interesting from the point of practical use because it provides spectral isolation between the two upper and lower bands of the same navigation signal. In this way it is possible to track each element separately or together.

### 5 Conclusions

The Power Spectral Density (PSD) for different types of BOC signal is shown. The analysis was carried out in this aspect comparing the BOC and CBOC modulations and then the method of generating the BOC signals was considered. In addition, the necessary tests and simulations of the process of generating CBOC (1.3), CBOC (12.1), CBOC (6.12) and CBOC (12.12) signals modulated with both sine and cosine subcarriers were performed using the Matlab/Simulink programming environment. The simulation tests were successful. Signal waveforms have been generated for channel B navigation data, B and C channel dispersing sequences and modulated B and C subcarrier channels.

The rectangular waveform of channel C dispersing sequence with primary and secondary code has a longer period than the dispersing sequence of channel B with only the basic code. The waveforms of modulated B and C subcarrier channels are essentially no different (Figures 1-2).

In general, graphical programming languages facilitate system design since they clearly define the relationship between the system modeling, simulation system and system implementation [35-36]. Since the Galileo system is still being developed and its improvements are expected, the *E*1 signal simulation shown in this work may contribute to facilitating the redesign of the transmitter and receiver.

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In addition, the simulation model may have educational purposes due to the transparency of the design and algorithm simulation. The introduced model may also facilitate system implementation. Implementation tools available in the Matlab/Simulink module, such as "Code Composer Studio", can be used at

the implementation stage.

On the other hand, since the transmitter design of most GNSS systems is similar, the simulation models introduced in this work can be used in design of the transmitter simulator, which will have a direct positive impact on the testing stages and checking the receiver.

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